

# Kick-off meeting GOCE+ Antarctica

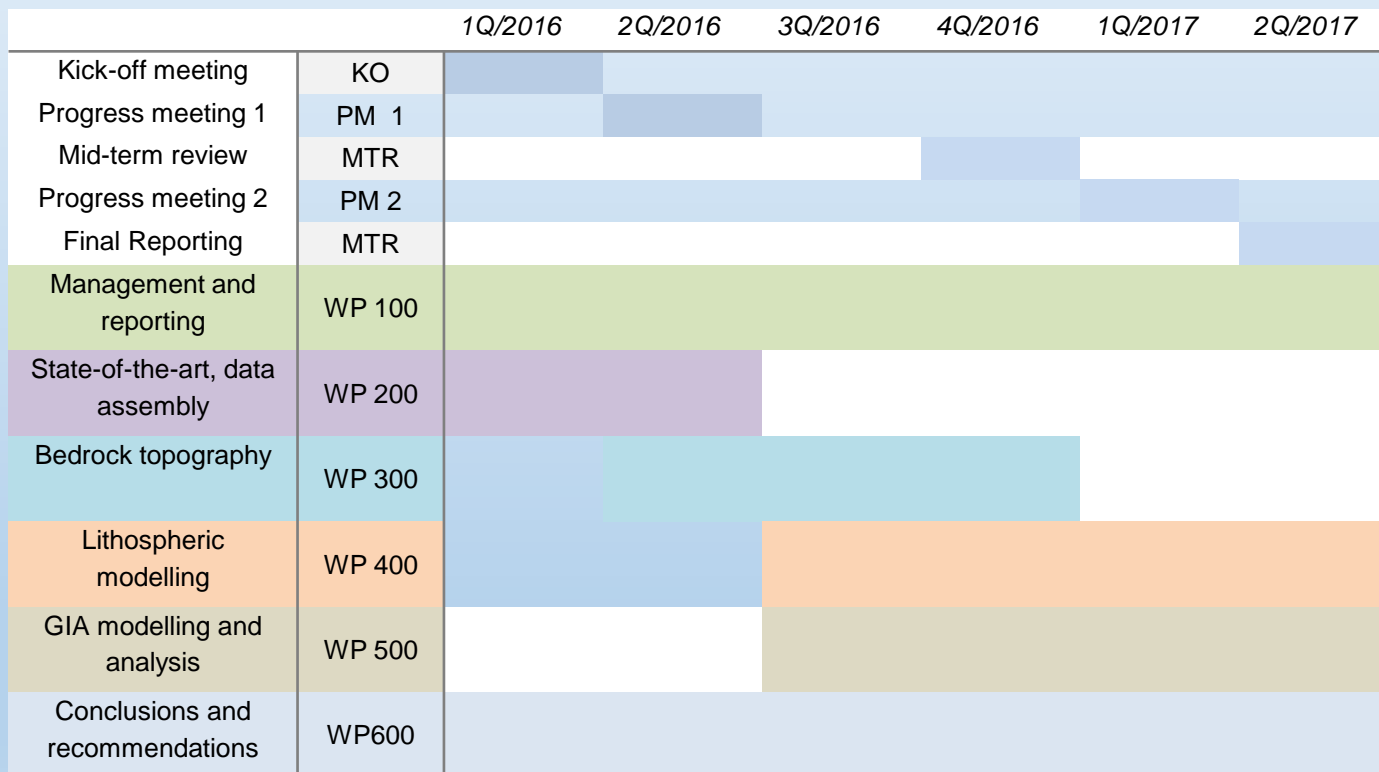
## **Dynamic Antarctic lithosphere**

Prepared by

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Ferraccioli (BAS), R. Forsberg (DTU Space)

# Timeline

Activity	Venue	Milestone	Target date
Kick-off meeting	ESTEC/Telecon (TBD)	KO	January 2016
Progress meeting 1	Kiel	PM1	May 2016
Mid-term review	ESTEC	MTR	October 2016
Progress meeting 2	BAS	PM2	February 2017
Final Reporting	ESTEC	FR	June 2017



# WP100 Management and reporting

- ESA-CAU contract is signed (18/01/2016)
- Sub-contacts with BAS, DTU, TU Delft
- Johannes Bouman is added as associated project member
- Monthly progress reports- first one sent

# WP 200 State-of-the-art, data assembly, modelling

Ice thickness

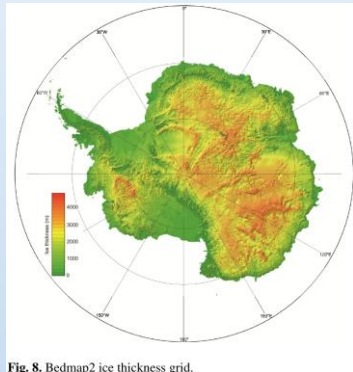


Fig. 8. Bedmap2 ice thickness grid.

Fretwell et al., 2013

(Sub-glacial) bed elevation

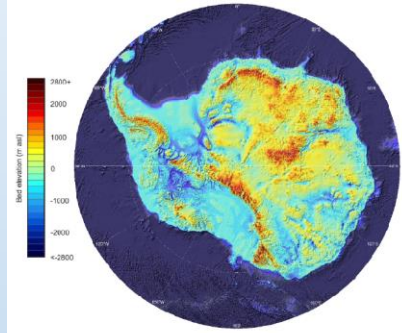
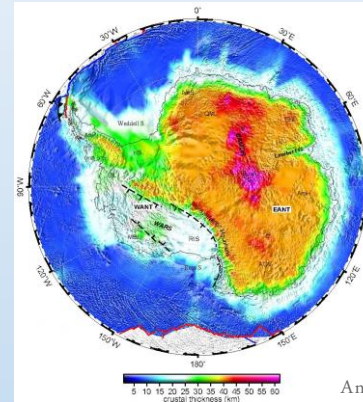


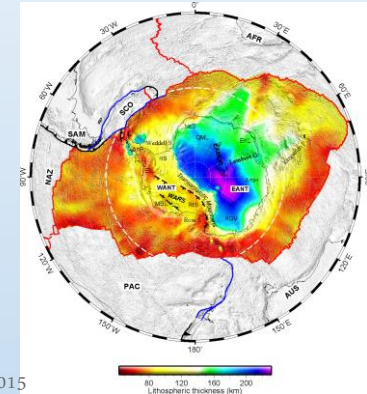
Fig. 9. Bedmap2 bed elevation grid. Although difficult to see at this scale, the bed elevation in areas where the main source of bed elevation data is gravimetric has inherited roughness from the surface grid.

Crustal thickness

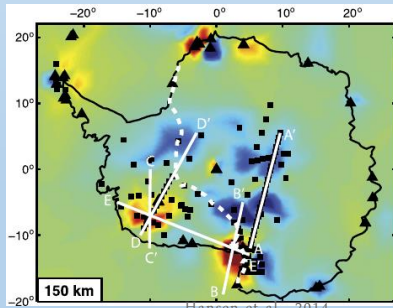


An et al., 2015

Lithospheric thickness

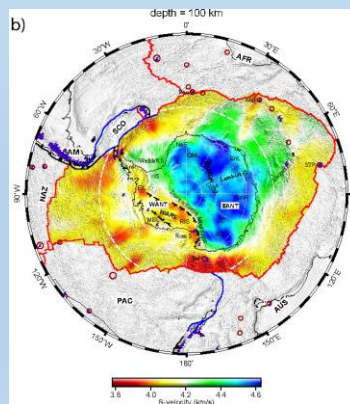


P-wave velocity

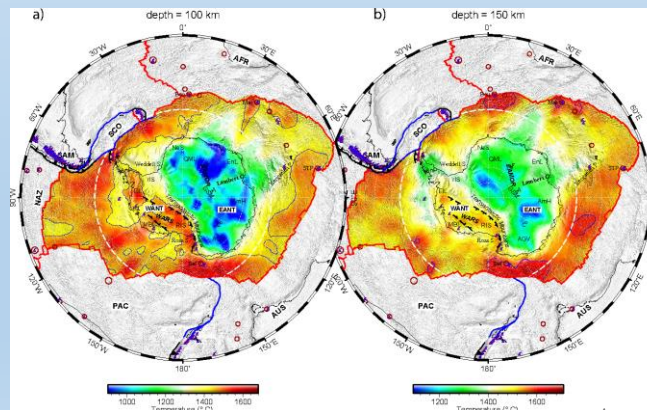


Hansen et al., 2014

S-wave velocity

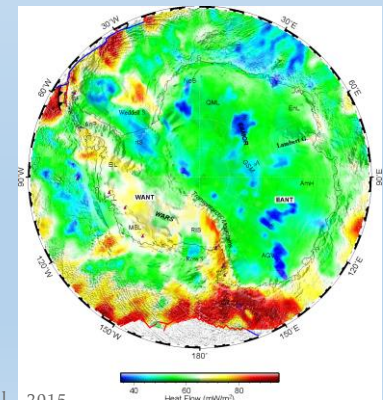


Temperature

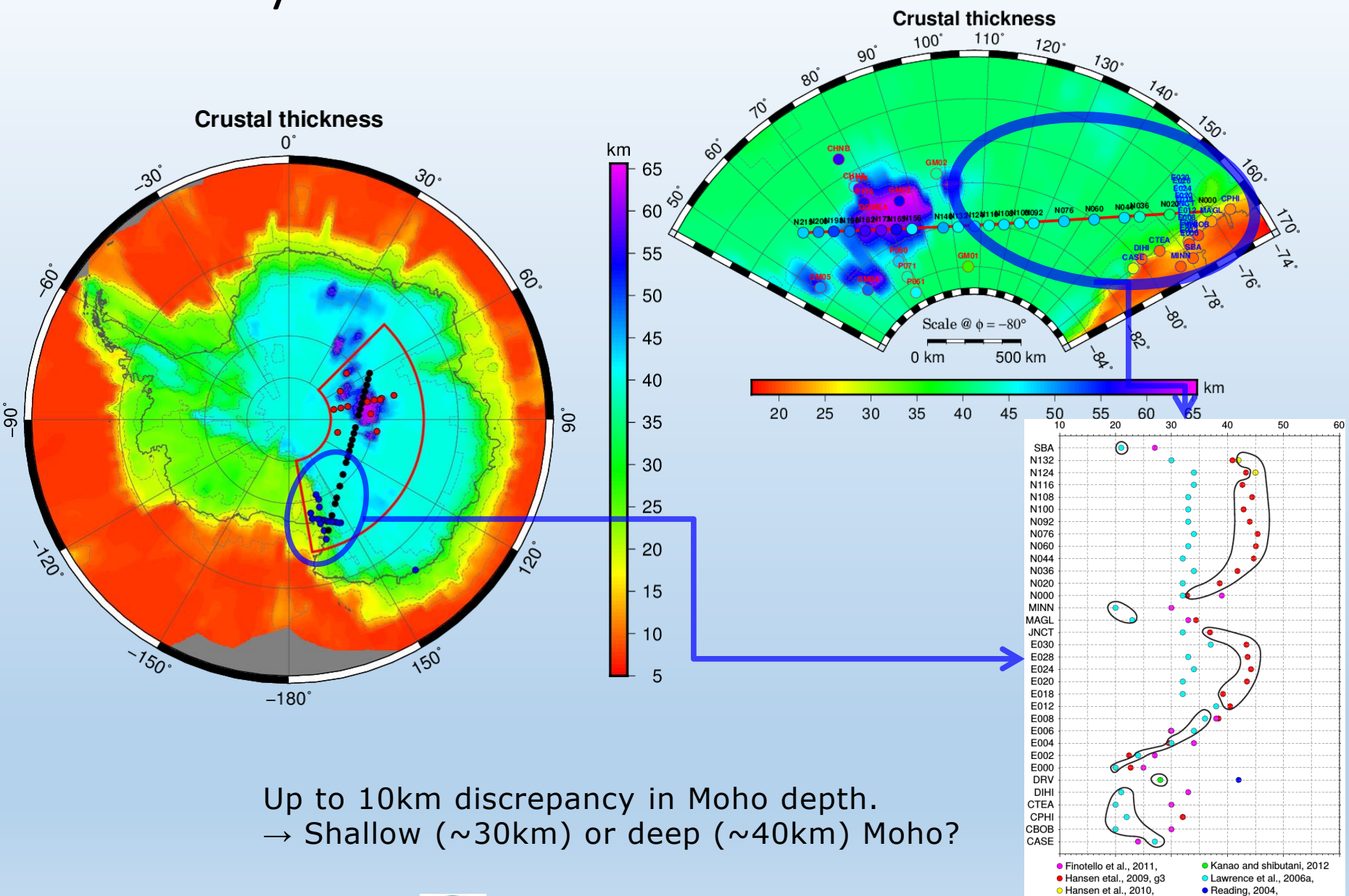


An et al., 2015

Heat flow



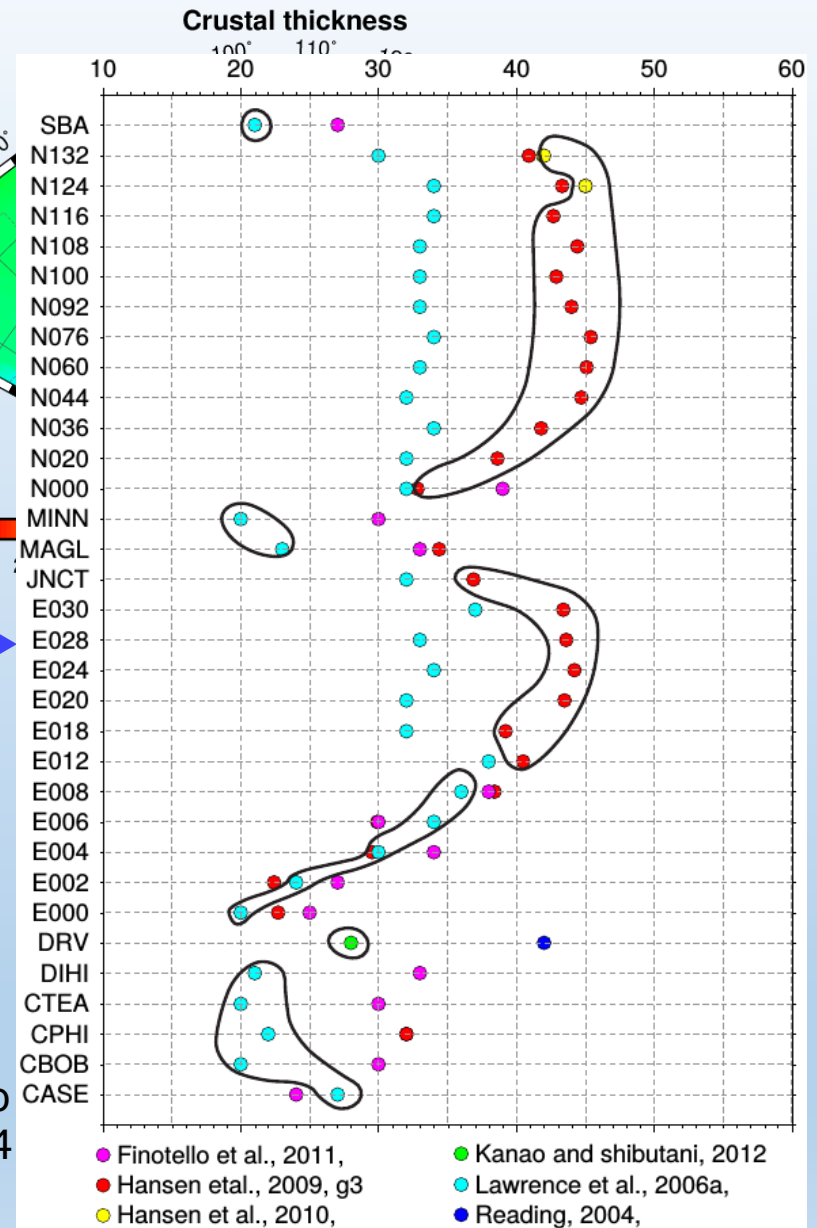
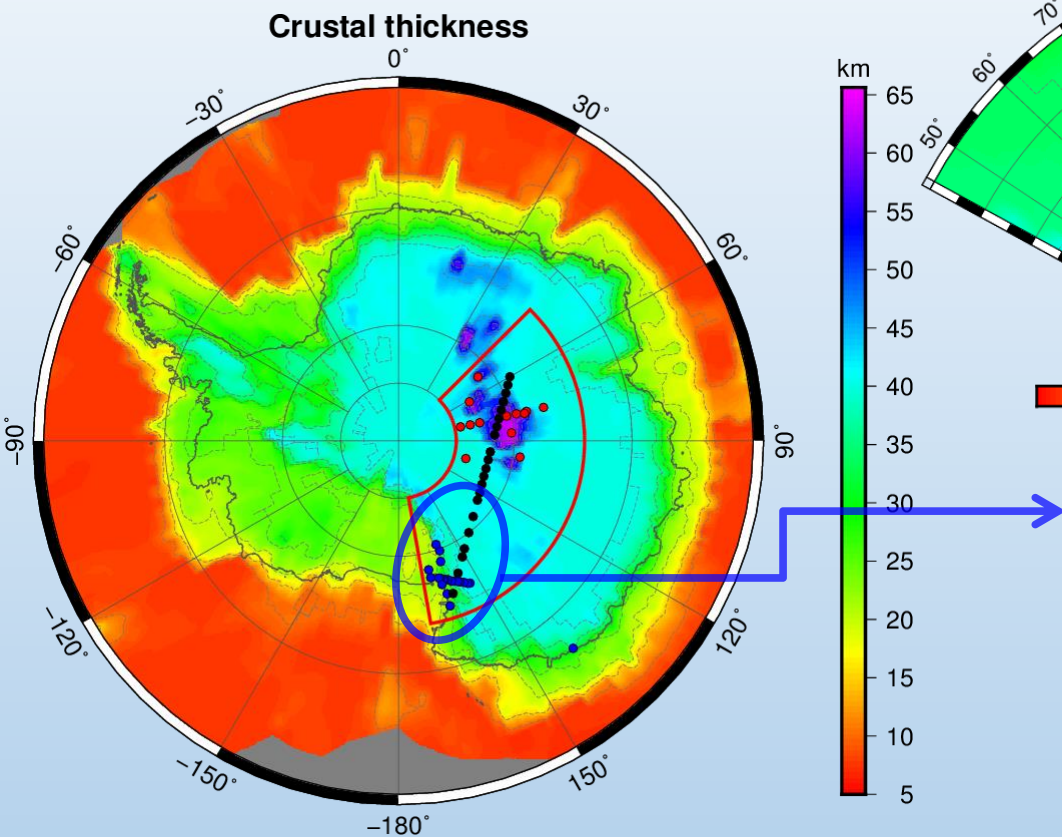
# Accuracy of seismic estimates



Up to 10km discrepancy in Moho depth.  
→ Shallow (~30km) or deep (~40km) Moho?

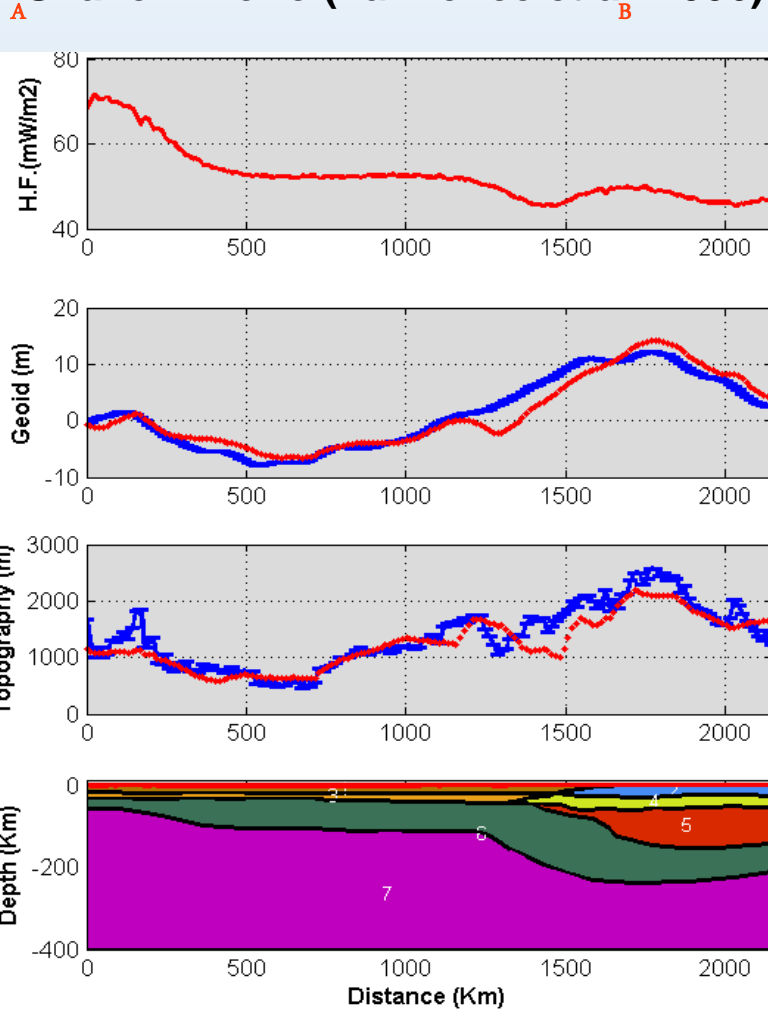


# Accuracy of seismic estimates

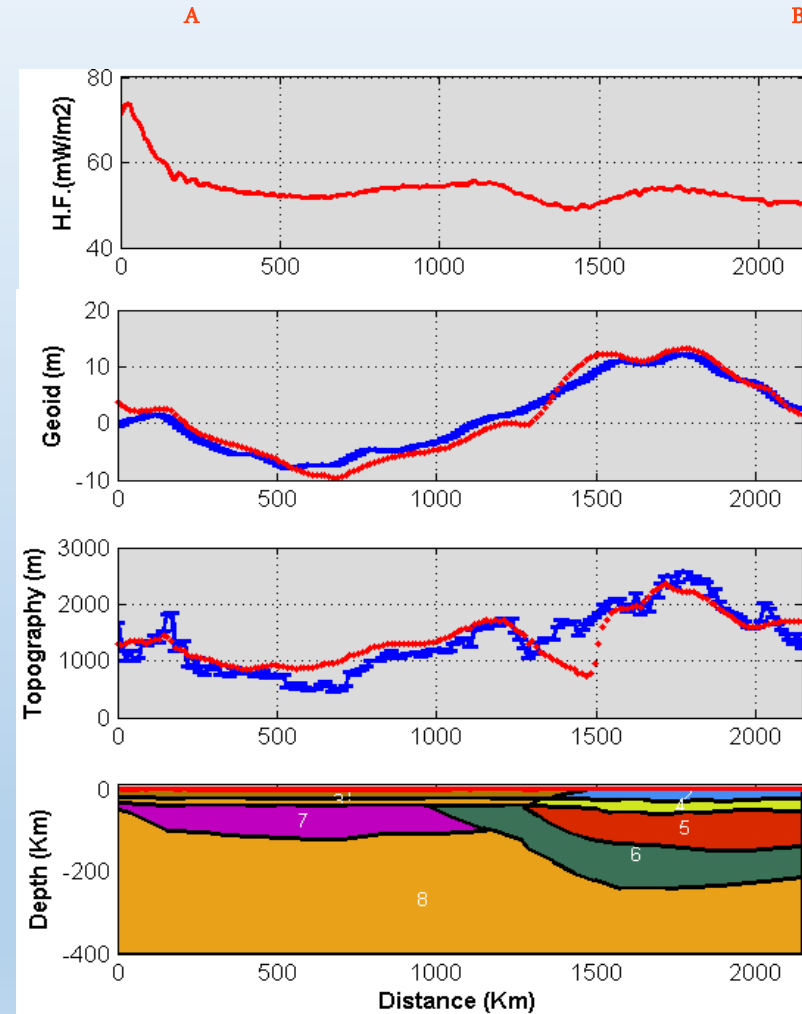


# Sensitivity to Moho depth estimates

**Shallow Moho (Lawrence et al. 2006)**

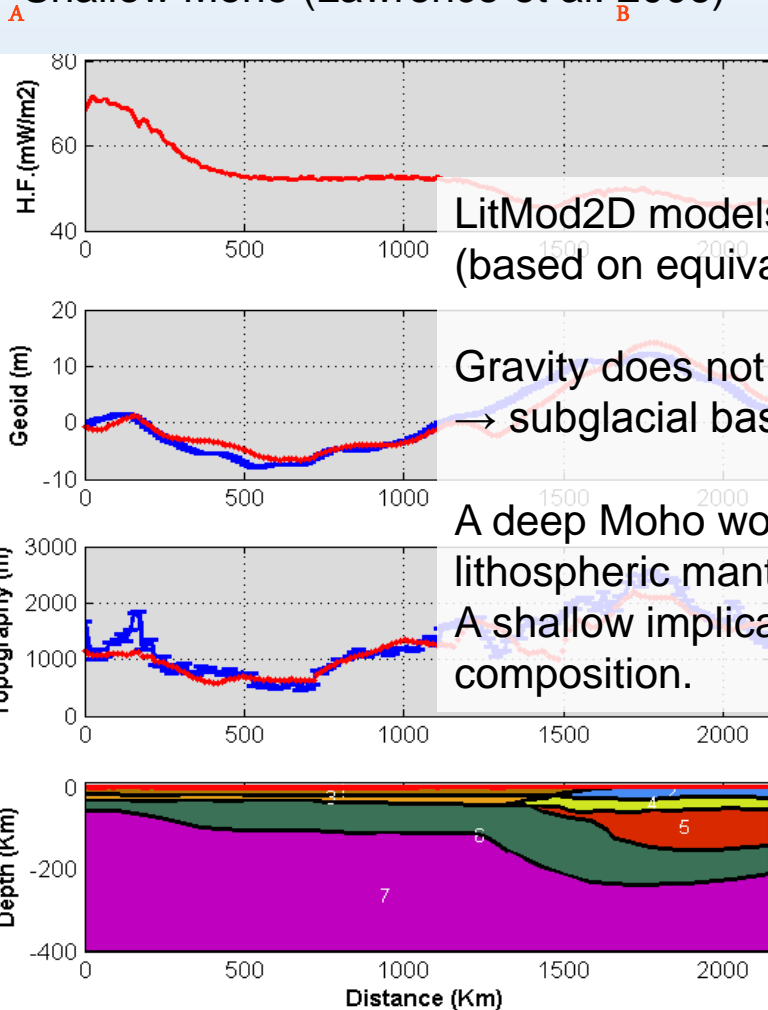


**Deep Moho (Hansen et al. 2009)**



# Sensitivity to Moho depth estimates

Shallow Moho (Lawrence et al. 2006)

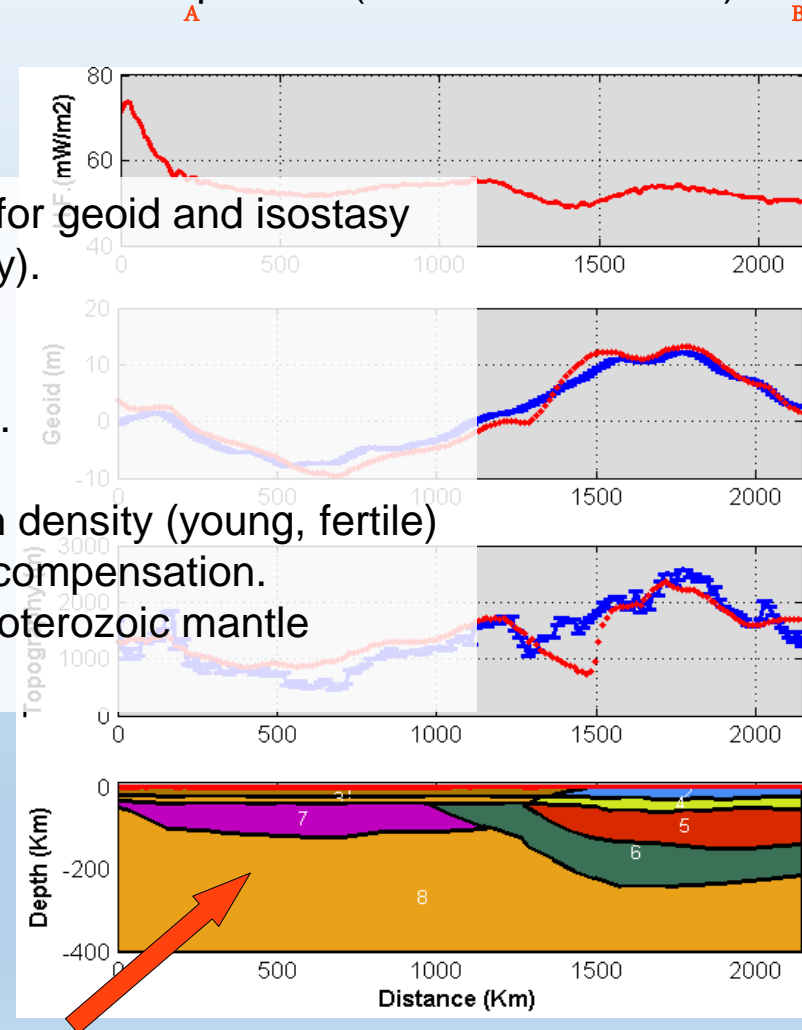


LitMod2D models fitted roughly for geoid and isostasy (based on equivalent topography).

Gravity does not fit  
→ subglacial basin not modeled.

A deep Moho would require high density (young, fertile) lithospheric mantle for isostatic compensation.  
A shallow implicates average Proterozoic mantle composition.

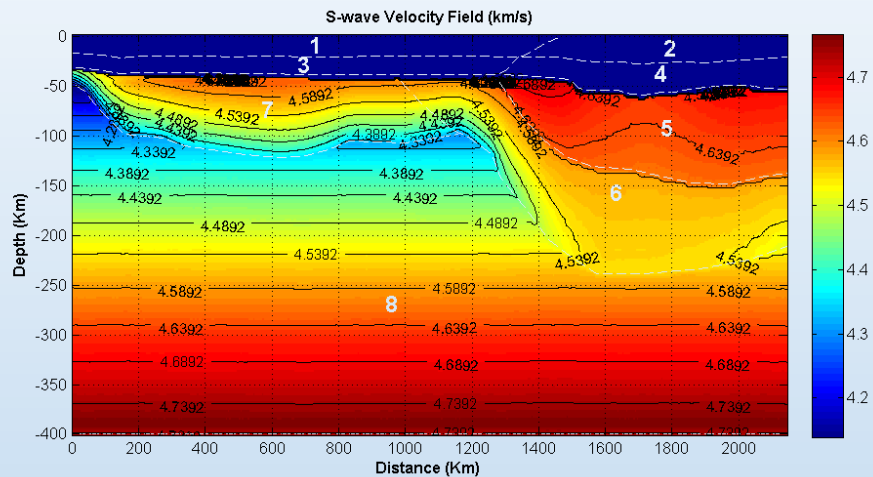
Deep Moho (Hansen et al. 2009)





# Comparison of S-wave velocities

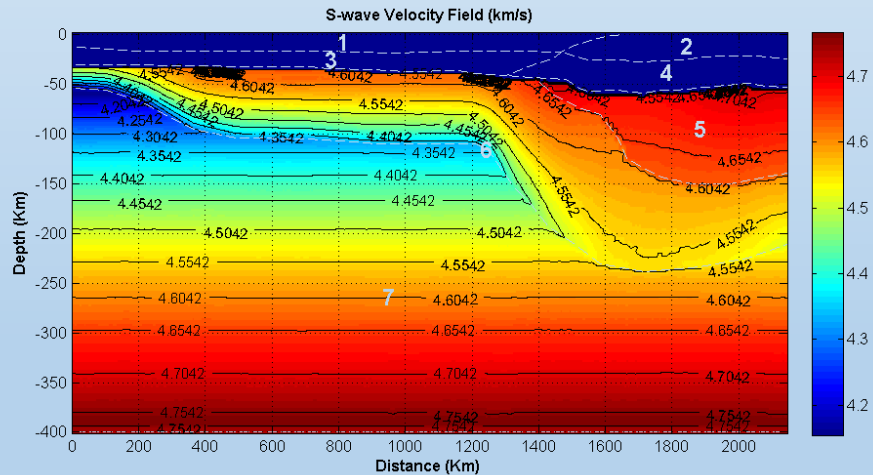
## LitMod model with deep Moho



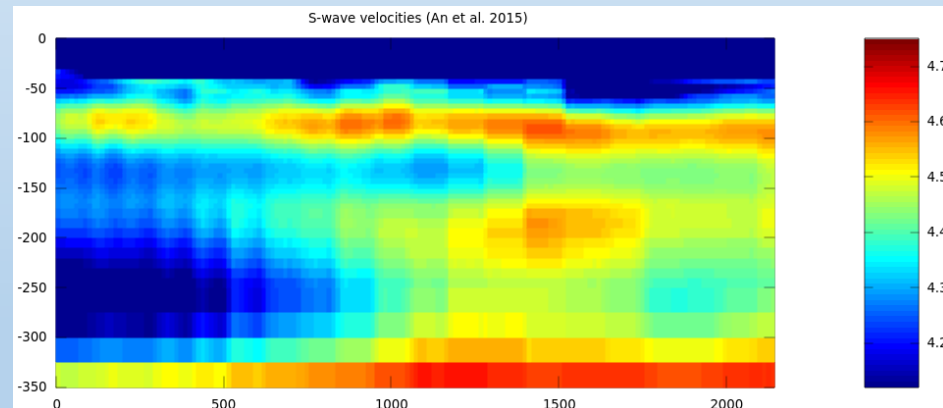
Fitting the models to topography and geoid results in reasonable velocity patterns. (However, there are still huge differences compared to the data.)

Neither a deep nor a shallow Moho setting can be yet excluded by the modelling of mantle velocities.

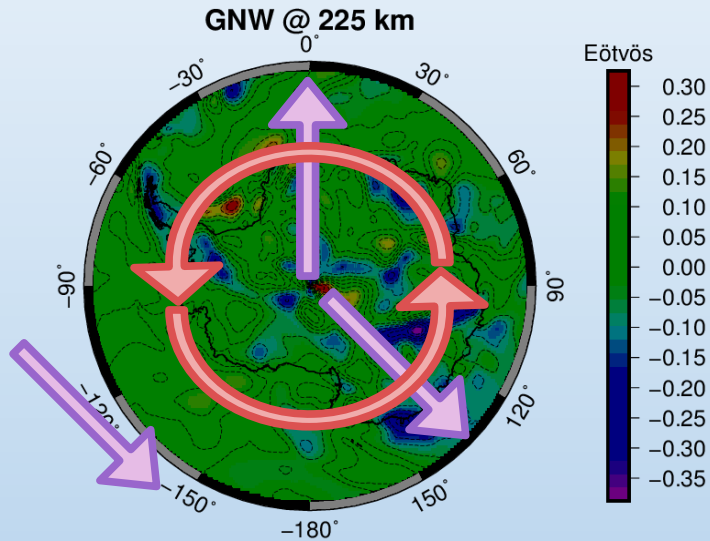
## LitMod model with shallow Moho



## Model “AN1-VS” from An et al. 2015

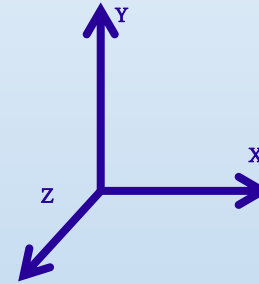


# Gravity gradient modelling



Local north-oriented frame  
“NWU” is not appropriate for intuitive interpretation of gravity gradients in polar regions.

We want:



in polar stereographic map view.

For the future, we refer to this by:  
“Indian Atlantic Up” (IAU).

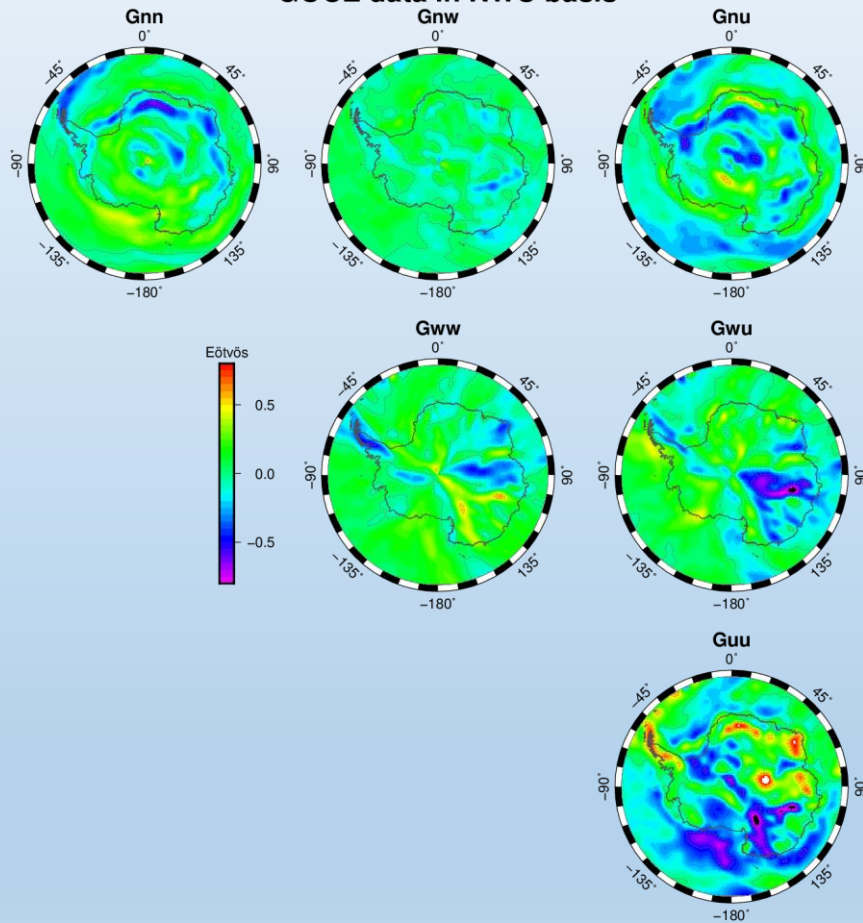
(Pointing to the Indian Ocean, the Atlantic Ocean and Upward.)

NWU (LNOF)

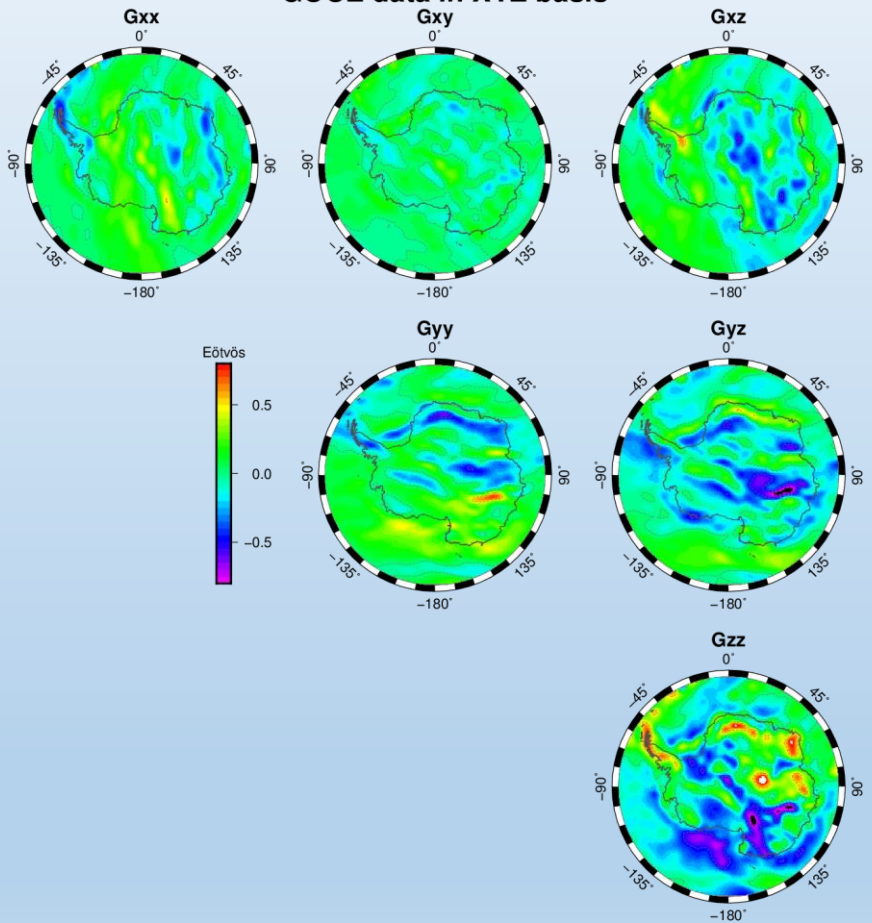
tensor basis change

IAU (to right, to top, up)

GOCE data in NWU basis

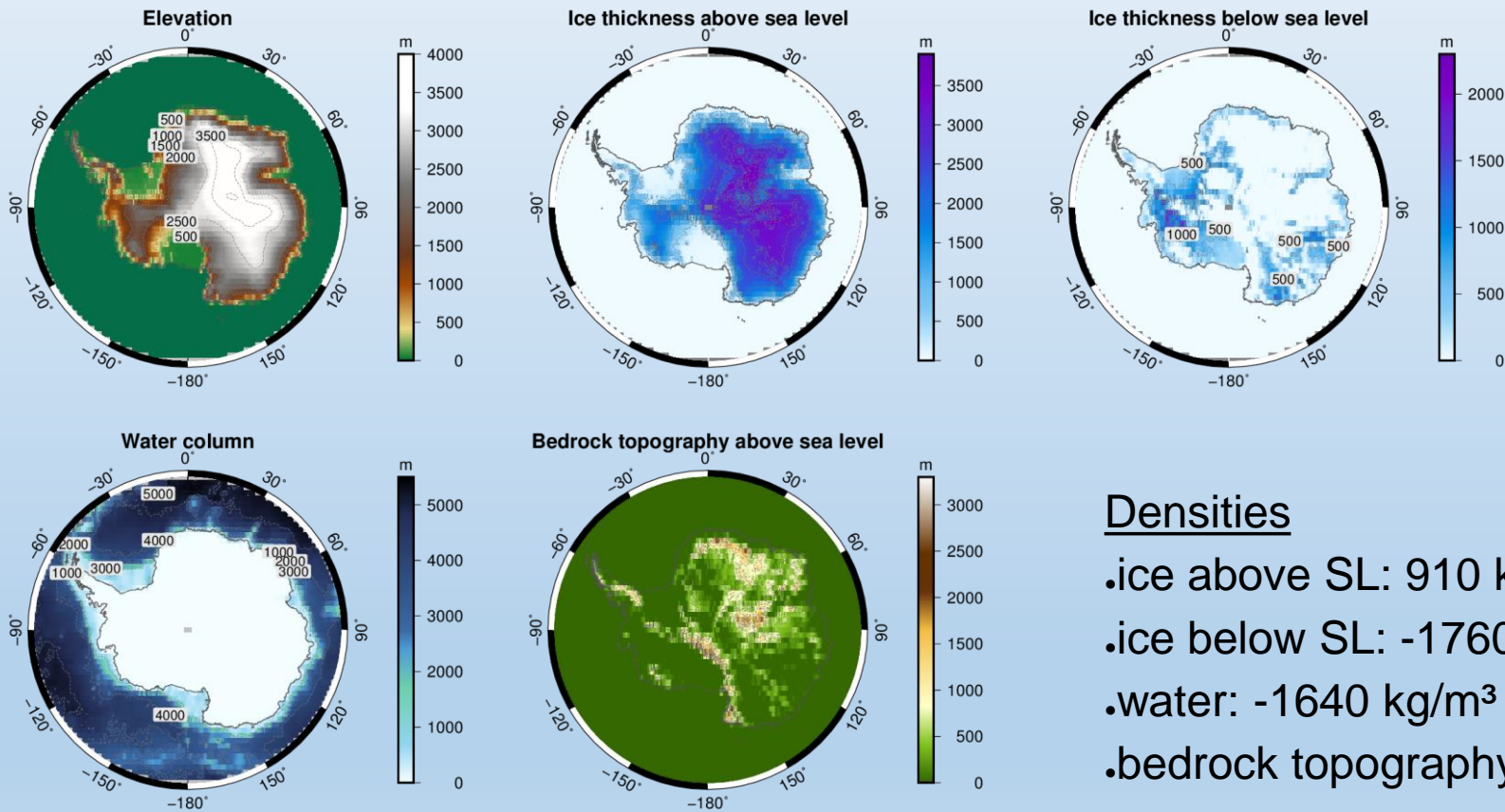


GOCE data in XYZ basis



# Topographic mass reduction

Calculation of gravity effect of ice, water and bedrock topography  
→ correction of GOCE's gradient tensor → intracrustal and subcrustal effects



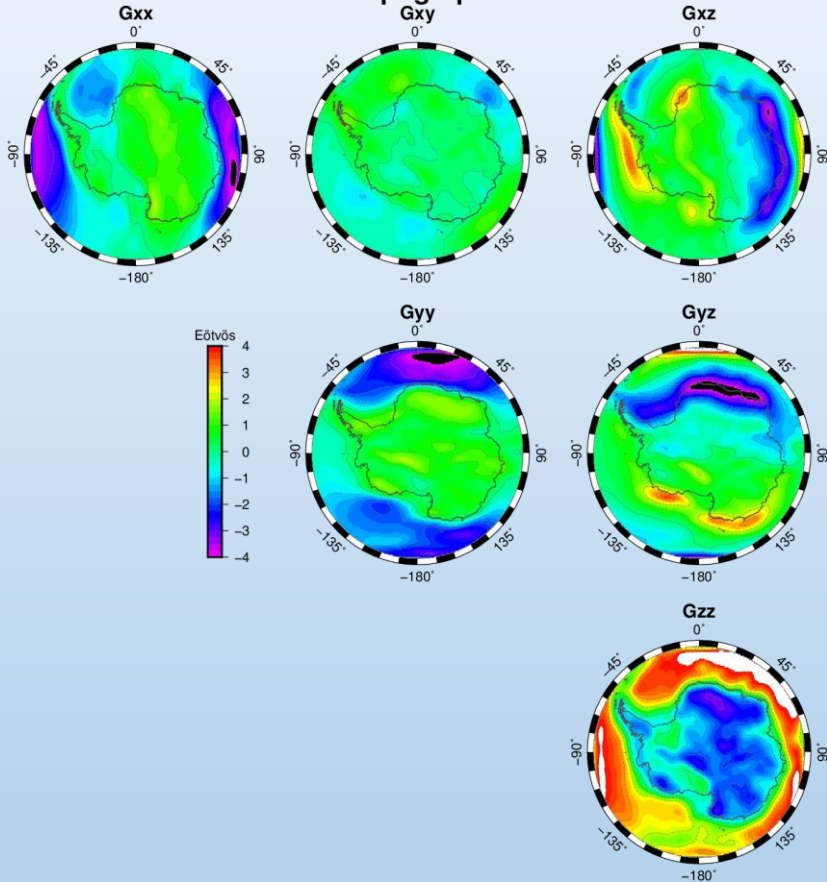
## Densities

- ice above SL:  $910 \text{ kg/m}^3$
- ice below SL:  $-1760 \text{ kg/m}^3$
- water:  $-1640 \text{ kg/m}^3$
- bedrock topography:  $2670 \text{ kg/m}^3$

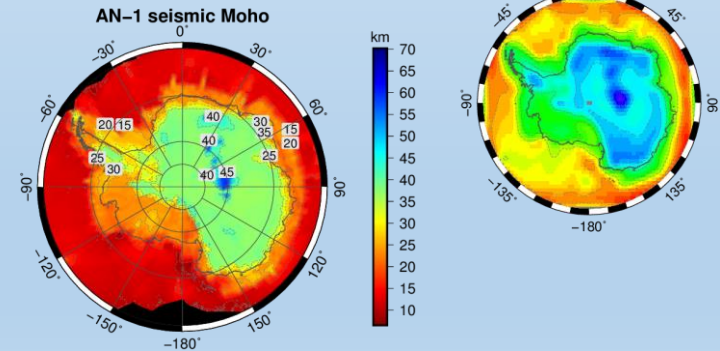
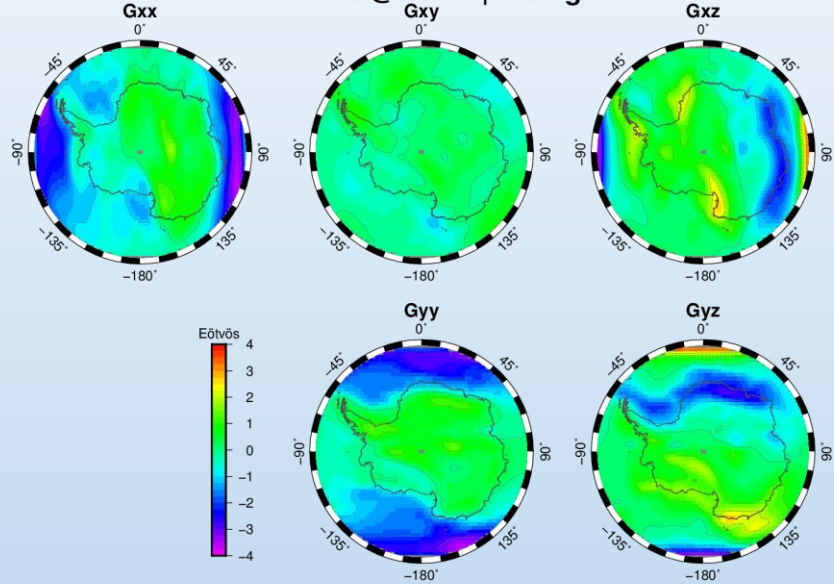


# Gradient signals from seismic Moho

GOCE after topographic correction



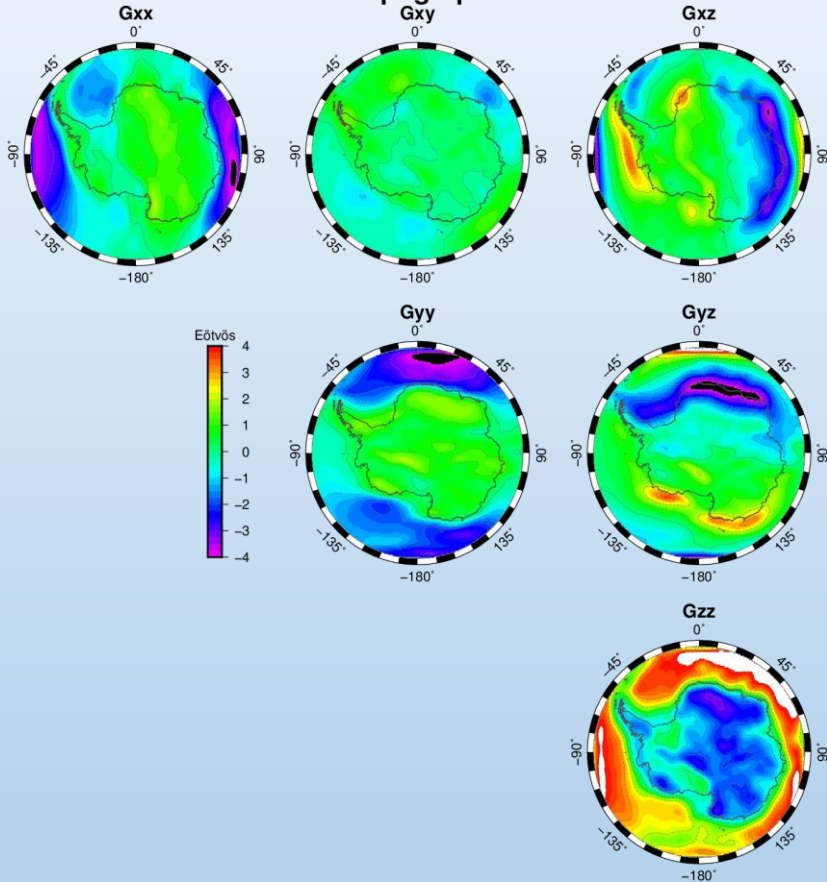
An-Moho @ 35km | 300 kg/m<sup>3</sup>



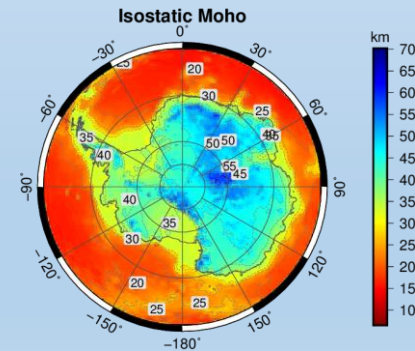
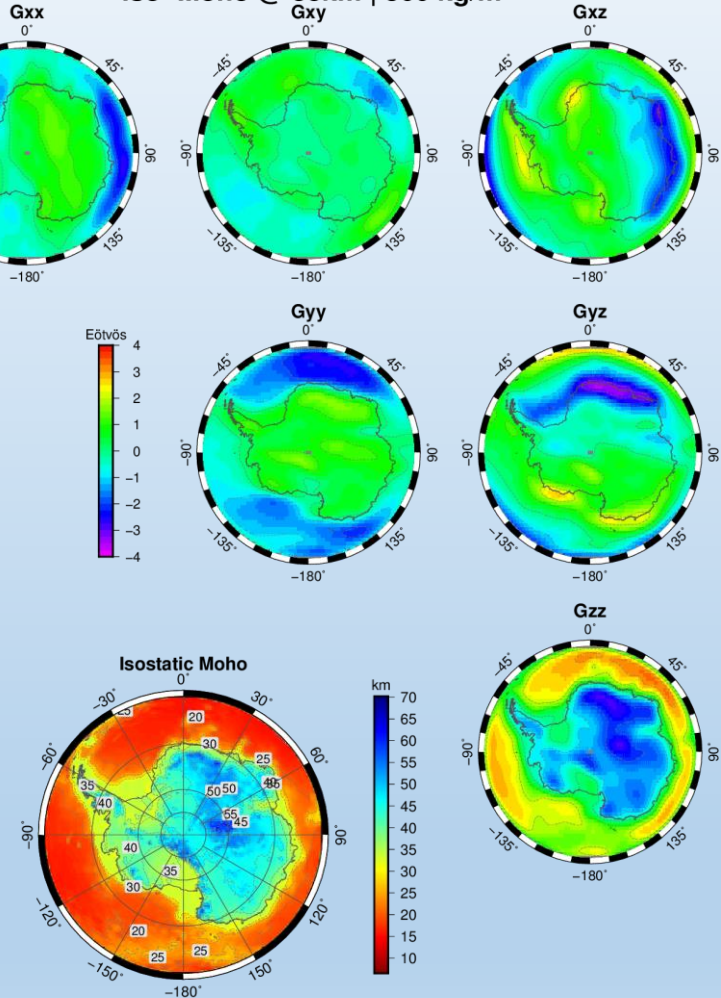


# Gradient signals from isostatic Moho

GOCE after topographic correction

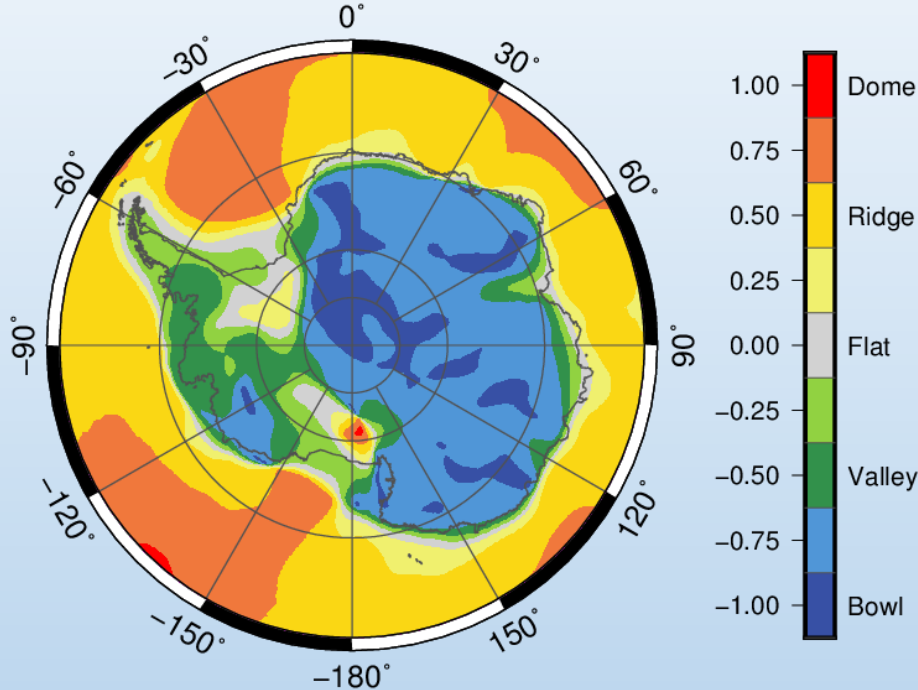


Iso-Moho @ 35km | 300 kg/m<sup>3</sup>



# Shape index of gravity gradients

## Shape index after topographic correction



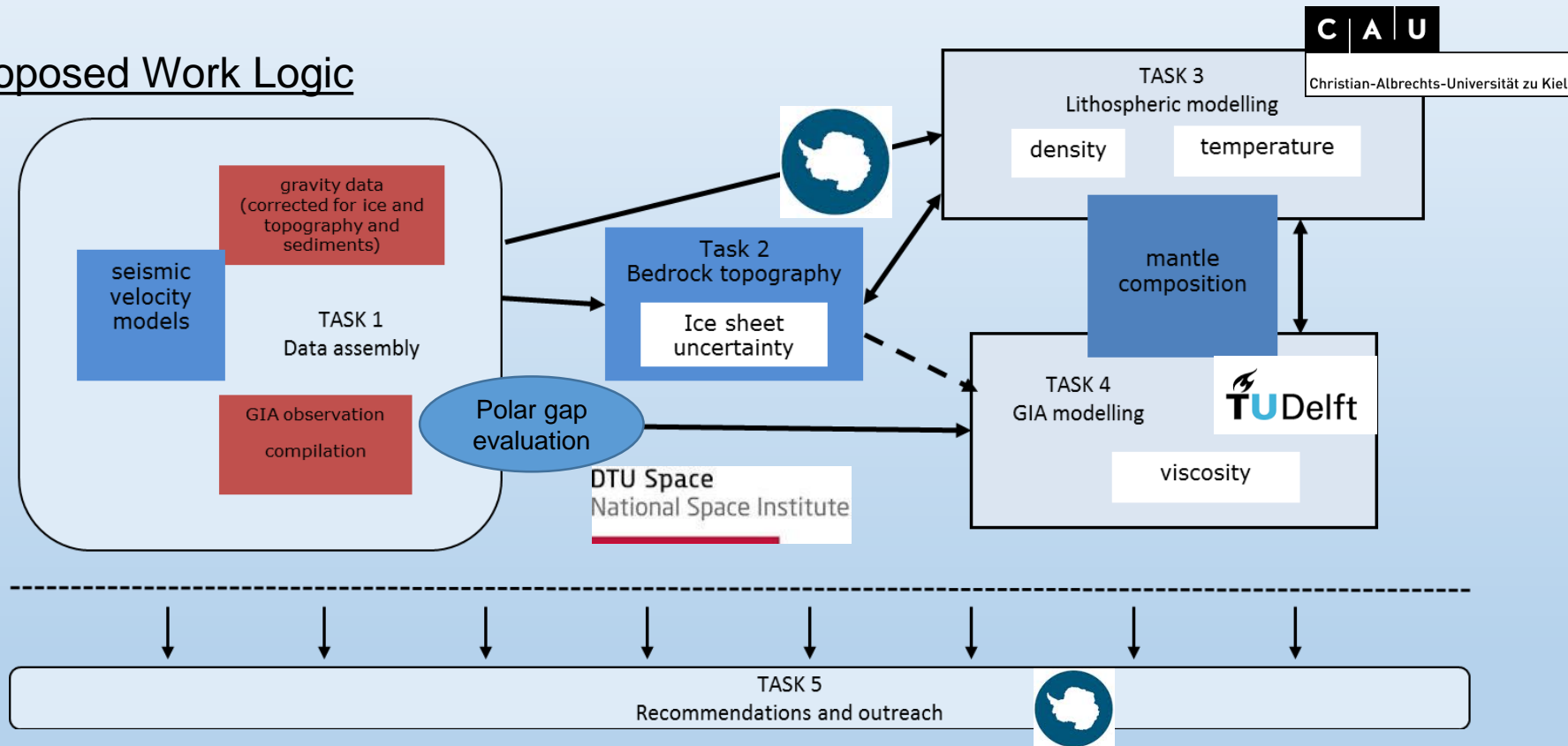
The shape index combines the minimum and maximum curvature of an equipotential surface (gravity potential) and can be expressed by the gradients.

Values and colours represent different morphologies of the equipotential surface, which would be difficult to identify in other images.

$$S_i = \frac{2}{\pi} \arctan \left( \frac{G_{zz}}{\sqrt{(G_{xx} - G_{yy})^2 + 4G_{xy}^2}} \right)$$

# WP 300-500 have not yet started

## Proposed Work Logic



# WP600 Outreach and presentations

- Abstracts to ESA Living Planet Symposium-submitted
- Abstract to DGG, Münster -submitted
- Abstracts to EGU 2016- submitted
- Abstracts for SCAR and IGC 2016- submitted
- Webpage